

Reconstruction of Horizontal Air Showers with AERA

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Surface Detector (SD) 1660 Water Cherenkov stations 1.5 km spacing 3000 km² covered area

Fluorescence Detector (FD)

27 telescopes at 4 sites with 180° view

Radio Detector (RD) → AERA

Auger Engineering Radio Array (AERA)



• Here: AERA124, ext. triggered

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- Build in phases
- Different antenna types and spacings
- 153 stations with 2 antenna (NS, EW) aera of ~17 km²
- Bandwidth
 30 80 MHz





Radio Emission from Cosmic Ray Air Showers



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Horizontal Air Shower (HAS)

- Full signal in RD, SD only sees muons
- Challenging reconstruction and MC
- Interesting features
 - \rightarrow e.g. v-search, composition, energy scale



Outline

- "large" radio footprints predicted, now measured
- Comparison of the signal in data & simulations
- Potential for sparse SD-RD hybrid array





Example Event

- Energy 19.6 EeV
- Zenith 79.2°
- Azimuth 351.7° (East)
- RD Stations 40



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 all antennas on one side of the shower axis

→ no asymmetry from the charge-excess contribution



Increasing Footprint

- RD externaly triggered by SD regular array
- 702 events in 20 month data (07/2013 02/2015)
- Footprintsize in shower plane increases for higher zenith angles





Simulations

• Similar reconstruction of simulations with real noise from the event time





AERA 1500m Array

- Only select stations on a 1500m grid
 - 10 ext. trig.
 - 3 int. trig.
- 51 Events

 found in 18
 month of data
 (03/2015 10/2016)
 4





Quick look into the data...



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Quick look into the data...



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Shower Reconstruction

- Accurate reconstruction of shower direction
- "problems" with energy estimation
 - Reconstruction designed for vertical showers
 - \rightarrow optimize
- Fully efficent for
 E > 10 EeV and Θ > 65°





Summary

- HAS: large radio footprints on ground
 - → grid spacing ~km sufficient
 - → Sparse RD grid works for high energy and zenith angle
- Simulated amplitudes agree well with data

