

Super-Preshowers

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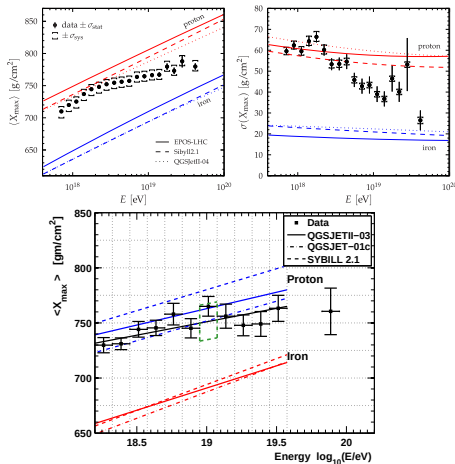


**PIERRE
AUGER**
OBSERVATORY

Motivation

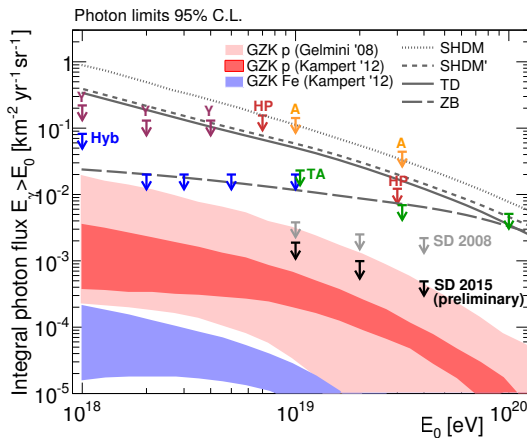
Observations at Pierre Auger:

- ▶ tension in Stereo and Hybrid data (Master thesis Markus Schauer)
dependent on $B_{\perp}$ (?) and E (like preshowers)
- ▶ composition puzzle:
 - ▶ $\sigma(X_{\max})$ shows heavier composition than $\langle X_{\max} \rangle$
 - ▶ inconsistencies between Auger and TA
 - ▶ muon excess (inconsistent with hadronic interaction models)

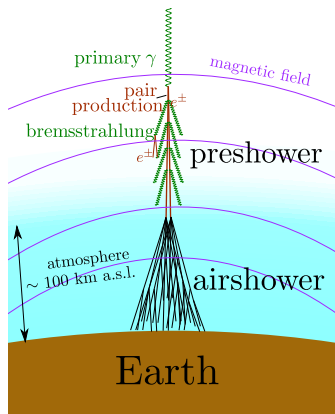


Motivation

- ▶ possible explanation: higher photon fraction + super-preshowers (SPS)?
- ▶ higher photons fractions imply exotic production models



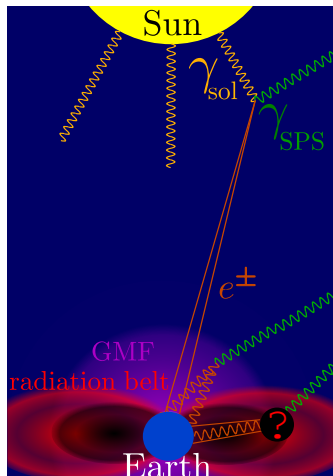
SPSs vs. preshowers



► preshower process:

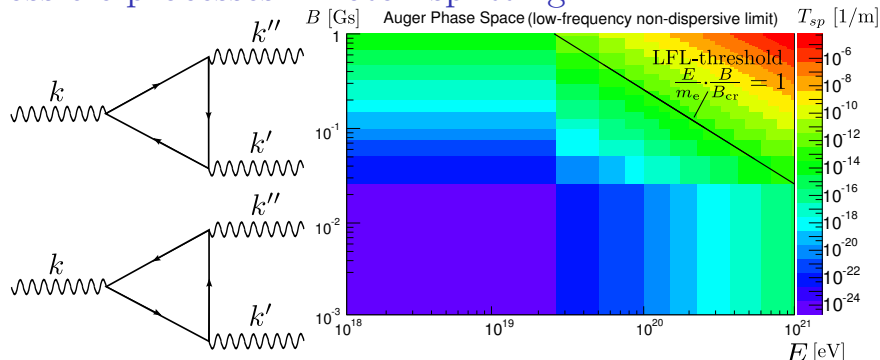
- pair production (PP) in geomagnetic field (GMF)
- bremsstrahlung by $e^\pm$ in GMF
- dependent on $B_\perp$ and E_γ

SPSs vs. preshowers



- ▶ SPSs occur at higher altitudes
→ PP in GMF too unlikely for SPSs!
- ▶ candidate mechanisms:
 - ▶ photon splitting in (geo-)magnetic field
 - ▶ pair production with solar photons
 - ▶ interactions in radiation belts (not yet studied)
- ▶ analysis:
 1. calculate probability of proposed mechanisms
 2. simulate SPS-induced air showers

Possible processes: Photon splitting



- ▶ photon splitting (PS) competing process to pair production (PP)
- ▶ PS can dominate PP for high B and low $E \rightarrow$ maybe also for Auger conditions?
- ▶ absorption coefficient too small in low-frequency limit (LFL)
- ▶ not clear for larger B and E (no convergence with literature outside LFL)

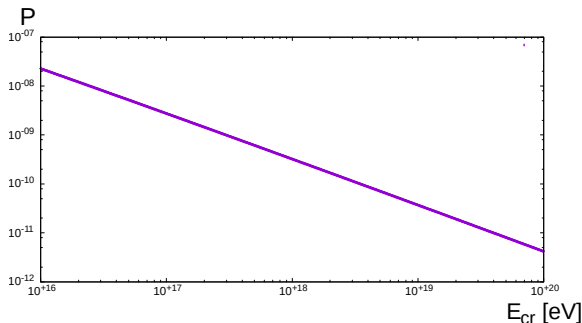
Possible processes: Pair production with solar photons

number density (from solar output/blackbody radiation)

$$dP = n \sigma dx$$

X-section (from literature on QED, e.g. Greiner)

thickness of layer



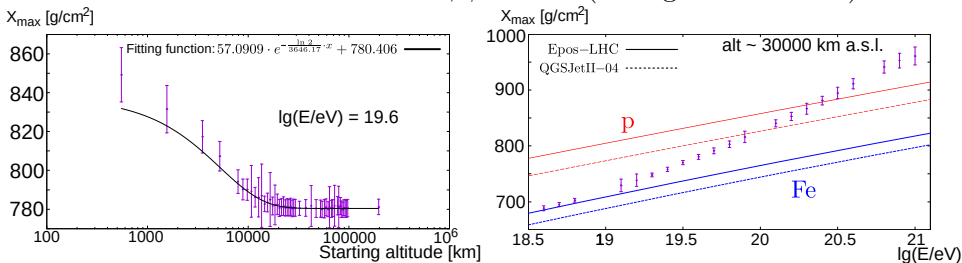
→ probability (after $x = 1$ AU) $\lesssim 10^{-8}$; too low!

Simulations with CONEX

- ▶ analyse effect of SPSs on observables
- ▶ here: energy deposit (X_{max} and longitudinal profile) and muon production (N_μ and dN_μ)
- ▶ implement SPSs in simulation code (PRESHOWER) by forcing first conversion

$X_{\max}$ vs. E

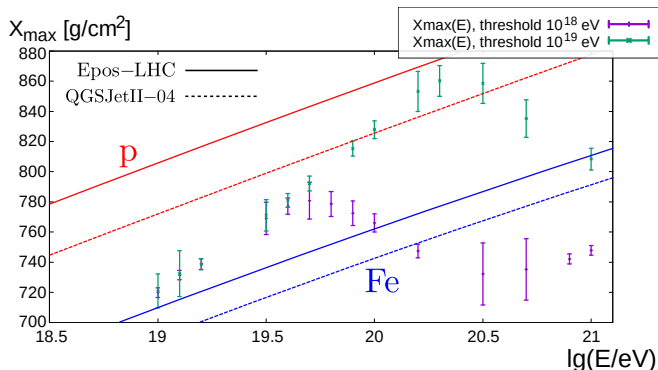
- ▶ $X_{\max}(\text{alt})$ and $X_{\max}(E)$
- ▶ arrival direction: $\theta = 60^\circ$, $\phi = 180^\circ$ (strong $B_\perp$ direction)



→ within hadronic range for large energy range
(above $\lg(E/\text{eV}) > 20.5$ no data!)

$X_{\max}$ vs. E

- ▶ secondary interactions: conversions for all photons energies above E_{thresh} (all at once!)
- ▶ shower configuration: strong $B_{\perp}$ direction

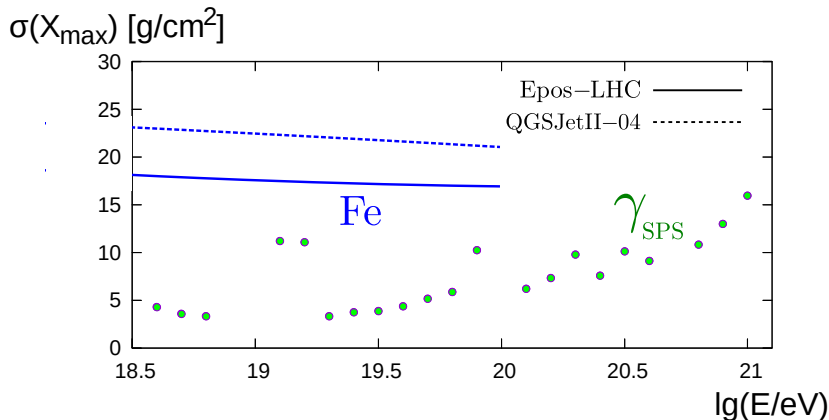


→ potential extension of acceptable range

→ details depend on processes (this is just “proof of concept”)

$\sigma(X_{\max})$ vs. E

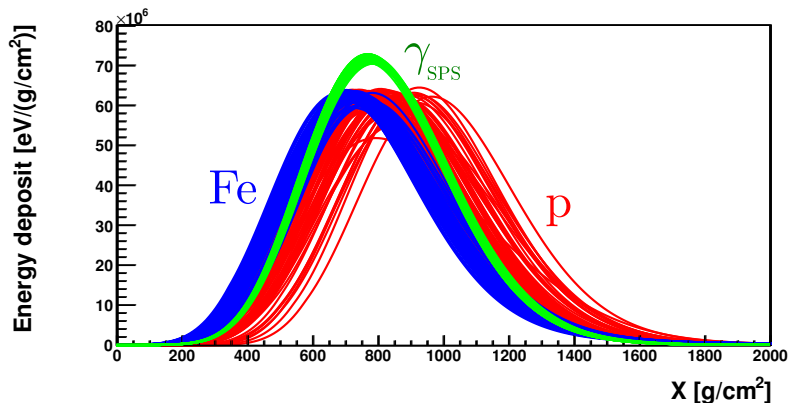
- ▶ shower configuration: strong $B_{\perp}$ direction; $\lg(E/\text{eV}) = 19.6$
- ▶ note: no error bars, “jumps” due to small number of showers



→ γ showers have less shower-to-shower fluctuations than hadrons!

Longitudinal profile of dE/dX

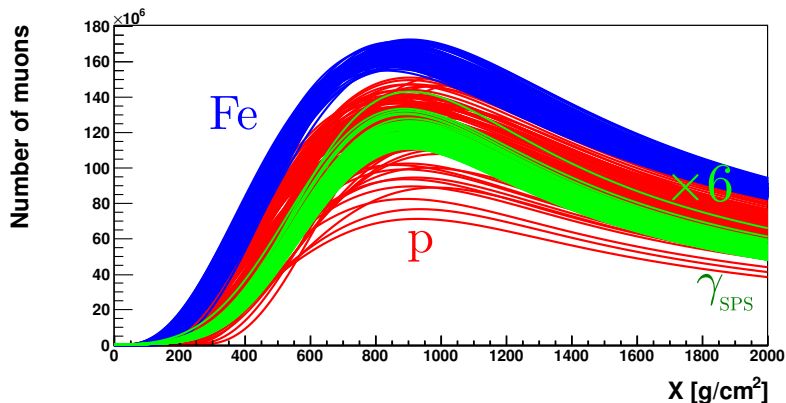
- ▶ shower configuration: strong $B_{\perp}$ direction; $\lg(E/\text{eV}) = 19.6$



→ longitudinal profiles also very similar (apart from shower-to-shower fluctuations)

Muon production

- ▶ shower configuration: strong $B_{\perp}$ direction; $\lg(E/\text{eV}) = 19.6$



→ lower muon production for SPS, but not so severe (maybe underestimation of energy? Photonuclear X-section?)

Summary and outlook

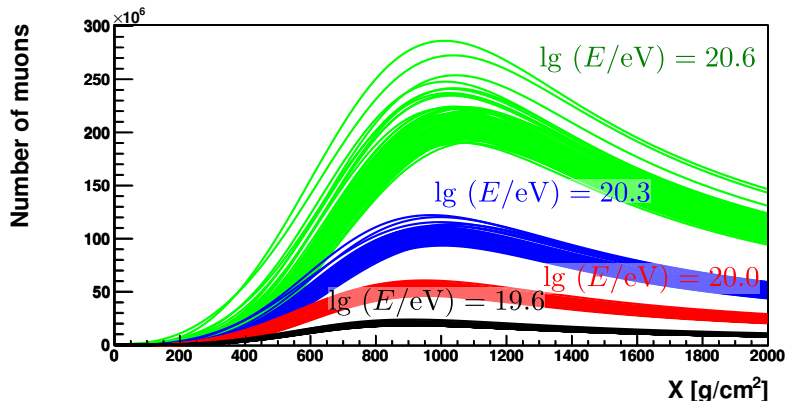
Summary:

- ▶ composition puzzle: not necessarily hadronic
 - ▶ SPSs could mimic hadronic showers for large energy range
 - ▶ SPSs may explain discrepancies and inconsistencies mentioned in motivation
- ▶ muon excess:
 - ▶ underestimation of energy for SPSs?
 - ▶ larger photonuclear X-section?
- ▶ no SPS process identified sofar! PP with solar photons ruled out!

Outlook:

- ▶ interactions in radiation belts a possible SPS mechanism?
- ▶ comparison for more shower observables (e.g. rise time)
- ▶ comparison with exotic production models (including E_{thresh} , other components, SPS simulations ...)

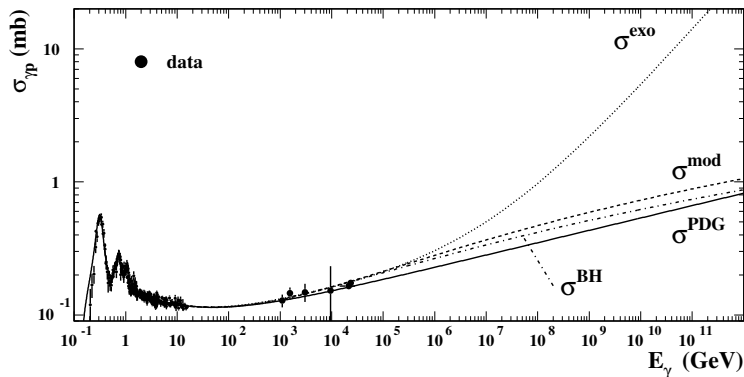
Muon production



- ▶ number of muons increases with energy

SPS simulations – N_μ

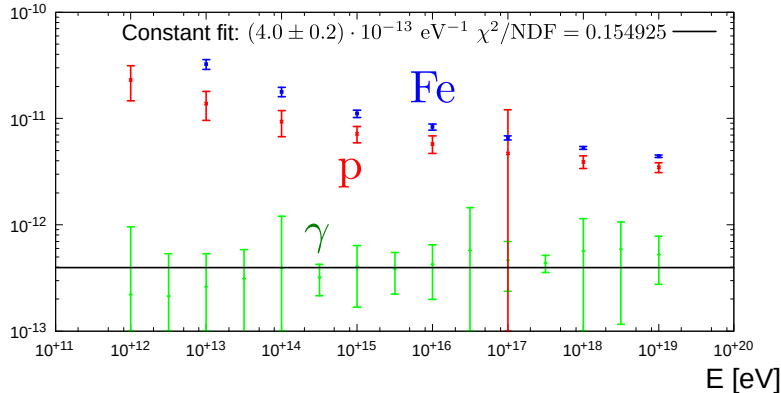
- ▶ photonuclear X-section extrapolated from lower energies
- ▶ higher values depending on model; normally, PDG trend used



Muon production

- ▶ not so consistent:

$N_\mu(X_{\max})/E$ [eV⁻¹]



- ▶ muon production X-sections not dependent on E for γ
- ▶ does this make sense for SPS (with many low-energy photons)?