

# Spherical Wavelets

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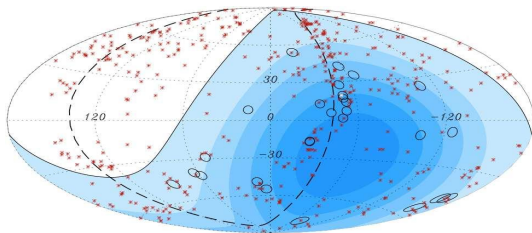
October 5, 2010

# Motivation

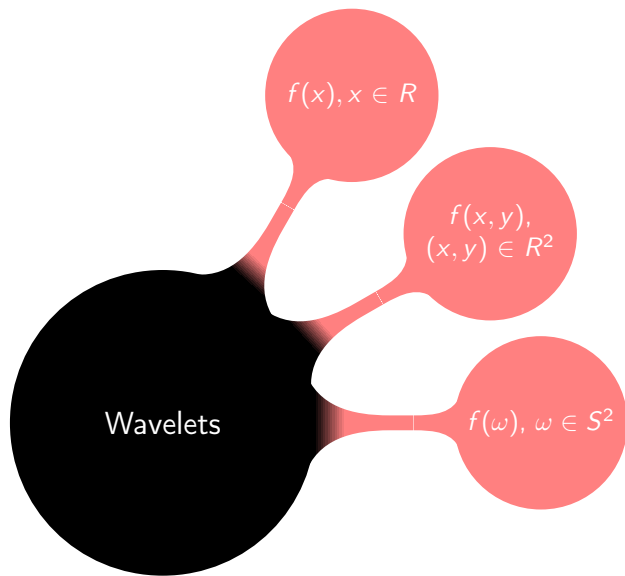
When are wavelets useful?

"... if we are interested in transient phenomena - a word pronounced at a particular time, an apple located in the left corner of an image - the Fourier transform becomes a cumbersome tool."(Mallat, Wavelet Tour)

In astrophysics we are interested on finding the sources(Pierre Auger) and on eliminating them(WMAP).

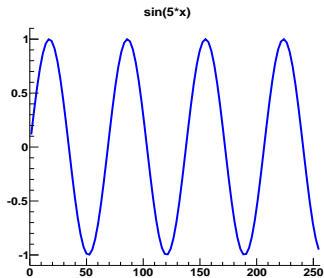
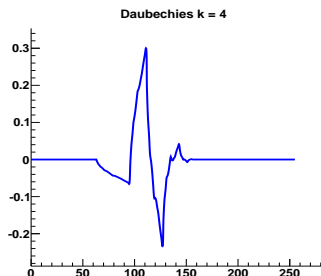


# What are wavelets?



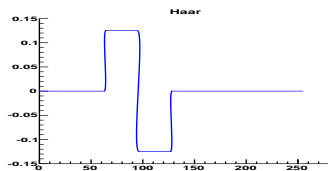
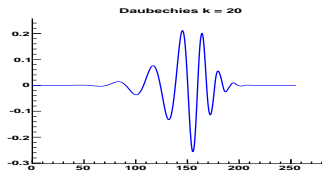
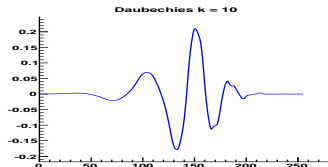
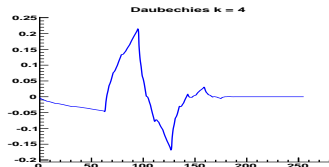
# Wavelets in one dimension

Functions with localization on frequency and time domain, whereas sines and cosines have localization only in frequency domain.



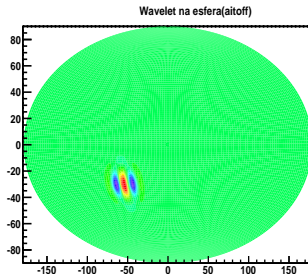
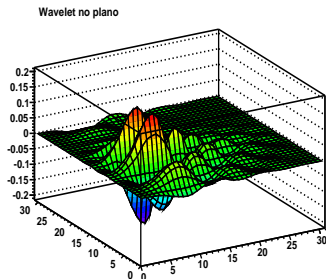
# Various possible wavelets to choose

## Some wavelets of the Daubechies family



# Wavelets in two dimensions

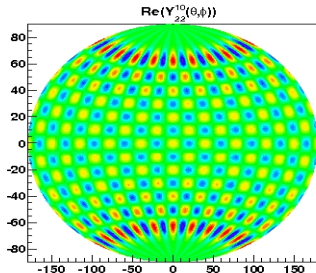
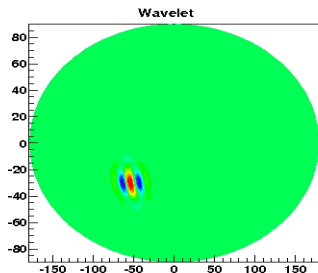
- Defined on the plane (useful for images analysis).
- Defined on the sphere (data collected in all directions, WMAP, Pierre Auger).



# Why wavelet and not harmonic analysis?

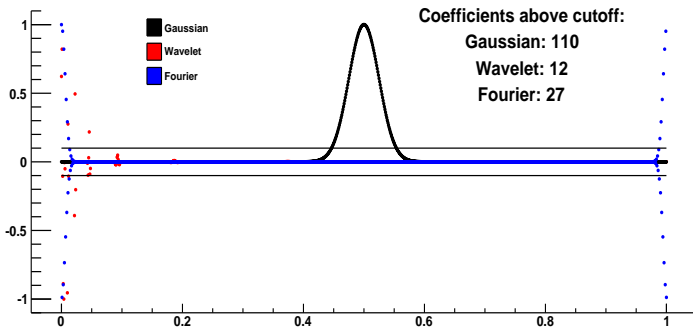
Cosmic ray sources will manifest themselves as point-like source in event maps (limited by angular resolution) or maybe stretched due to the deflection of their trajectories by galactic and inter-galactic magnetic fields, so we need local functions to look for them.

- Wavelets: Local.
- Spherical harmonics: Global



# Fourier, wavelet or coordinate domain?

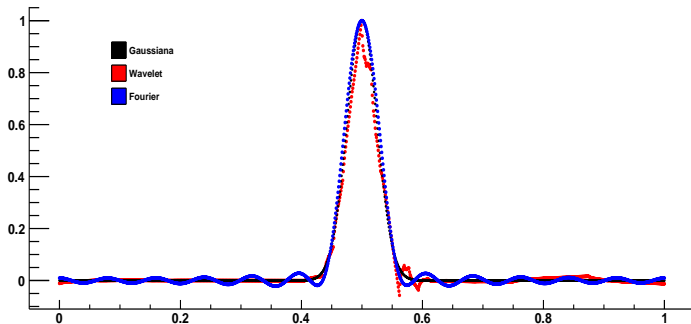
Representation of a Gaussian in Wavelet, Fourier and coordinate domain.  
We need less coefficients in wavelet domain to represent it!





# Reconstruction

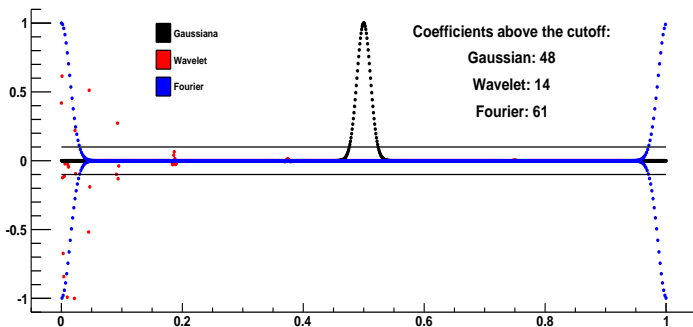
Wavelet reconstruction is more precise, since less coefficients are needed to represent it.



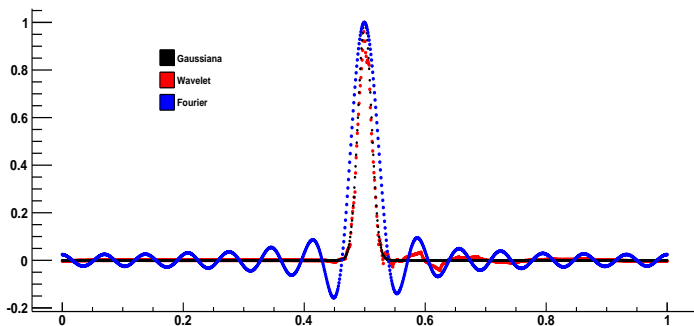
As the Gaussian gets narrower ...

# Fourier, wavelet or coordinate domain?

We need four times more Fourier coefficients than wavelet coefficients.



Fourier reconstruction becomes really bad.



- Coordinate domain  $O(B^5)$ .

$$C(g) = \int_{S^2} f(\omega) \overline{\Lambda(g)h(\omega)} d\omega$$

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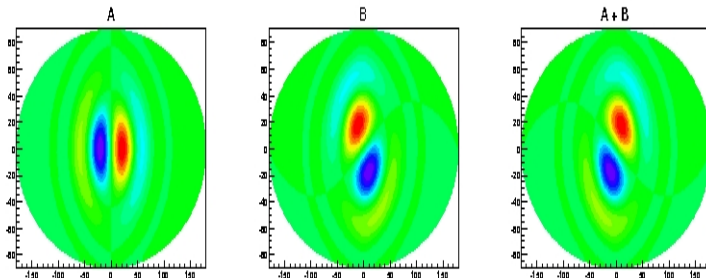
$$C(g) = \int_{S^2} f(\omega) \overline{\Lambda(g)h(\omega)} d\omega$$

- Harmonic domain  $O(B^4)$ .

$$W(\alpha, \beta, \gamma) = \sum_{l=0}^{B-1} \sum_{m=-l}^l \sum_{n=-l}^l \sum_{u=-l}^l \overline{a_{lm}} b_{ln} B_{mnu}^l e^{im\alpha} e^{iu\beta} e^{in\gamma}$$

# $O(B^4) \rightarrow O(B^3)$ : Steerable Wavelets

Using steerable wavelets the complexity is reduced to  $O(B^3)$  with ( $N \ll B$ ).  $N$  related to the wavelet band limit.



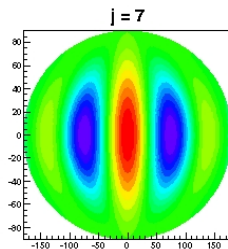
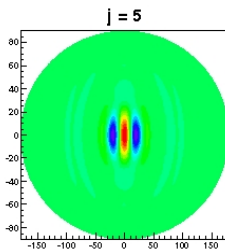
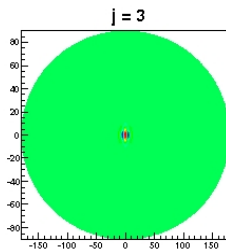
B	Package	64	128	256	512	1024
$\mu$ (MB)	S2DW	1.3	5.3	21	84	340
	SWAT	2	20	120	420	1600
$\tau$ (min)	S2DW	0.019	0.092	0.73	7	72
	SWAT	0.011	0.083	0.68	5.56	47.5

- SWAT measured on a 2.83 GHz Intel Core 2 Quad CPU.
- S2DW measured on a 2.20 GHz Intel Core 2 Duo CPU.

Wavelet is split into a kernel and a directional part. Dilations do not affect directional properties.

$$a_{lm} = K(l)S_{lm} \rightarrow K(2^j l)S_{lm}.$$

Example with  $J = 8$ :

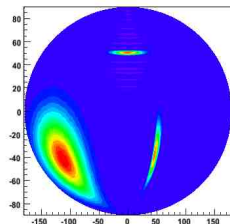




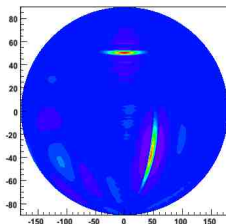
# Removing structures using scale decomposition.

Without the knowledge of its location, we can remove structures decomposing the map in scales. The map in the middle is a reconstruction using the 5 lowest scales out of 7.

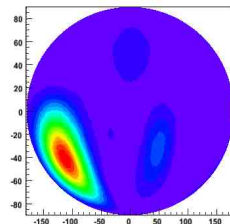
Original source



Decomposition using lowest scales.



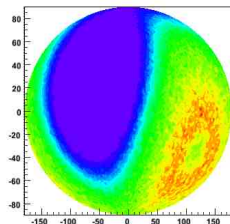
Decomposition using highest scales.



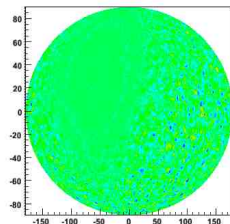
# Removing structures using scale decomposition.

Example on Pierre Auger data: Managing exposure effects. The sum of the two graphs on the right is exactly the map on the left.

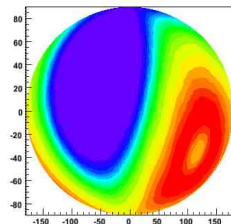
Events map



Lowest scales

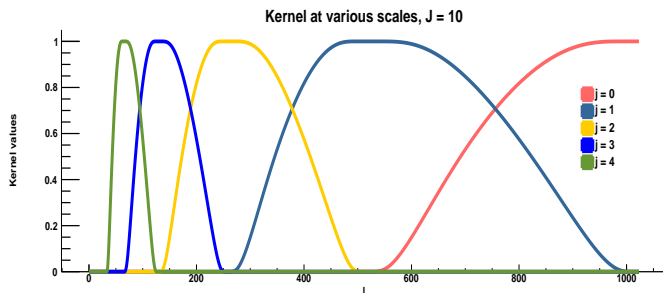


Highest scales



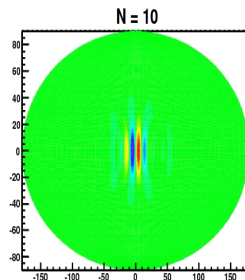
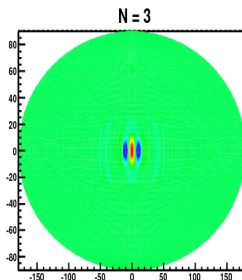
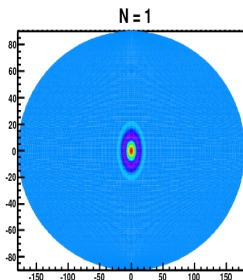
# Kernel has compact support

Only some frequencies are considered:



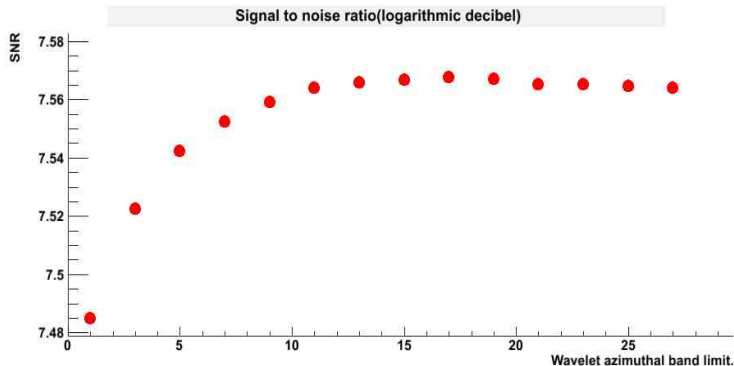
# Directionality control

Controlling the azimuthal band limit we can control directional properties.



# Application: increase SNR.

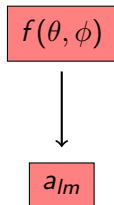
We can achieve higher signal to noise ratios if the wavelet correlates better with the source.



$$f(\theta, \phi)$$

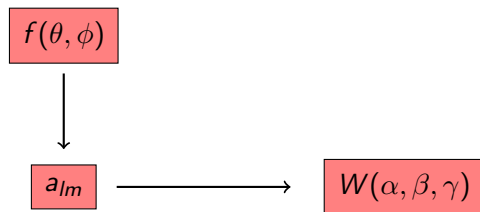
- 1 From Pierre Auger data, we calculate a map(Healpix).

# Denoising procedure



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- 2 Healpix is used to calculate SH.

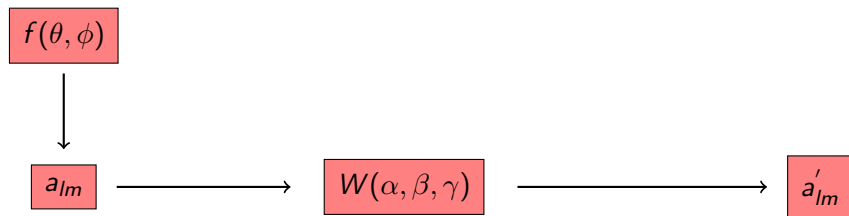
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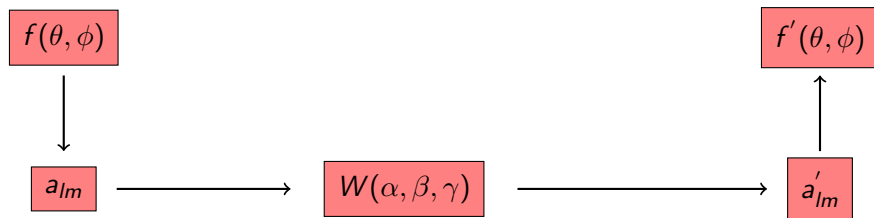


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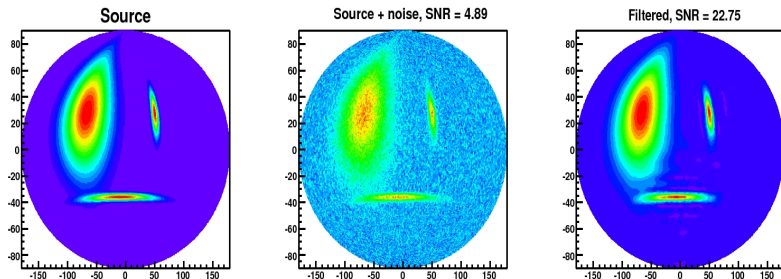


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Donoho and Johnstone threshold value:

$$T = \sigma \sqrt{2 \log M}, \quad \text{SNR} = 20 \log_{10} \left( \frac{\sigma_s}{\sigma_n} \right)$$

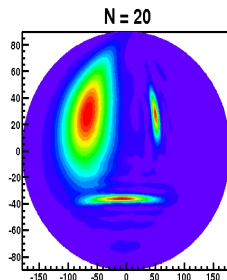
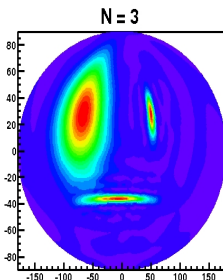
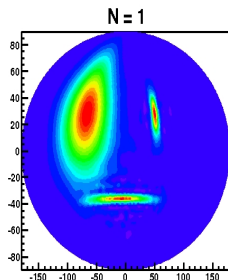
Where  $\sigma$  is the noise RMS and  $M$  the number of points. Denoising Gaussian beams immersed in a Gaussian noise:



# Signal to noise ratio (SNR)

For different values of  $N$  we can achieve higher SNR depending on the structures on the map:

	$N = 1$	$N = 3$	$N = 10$	$N = 20$
SNR	21.99	23.91	25.13	25.23



## Wavelets:

- Another tool on "spherical data" analysis.
- Enhance sparsity.
- Useful for denoising, calcule of anisotropy, etc.

## Code:

- Integrated with ROOT.
- Objects can be saved to root files.
- Hard work done by FFTW.
- Easily loaded in Cint(good!).
- Healpix code included in the build system.

Available at: <http://www.ifi.unicamp.br/~mzimbres>