Radiodetection of extensive air showers at the South Pole

Tobias Degner

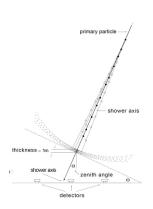
RFWU Bonn

Astroteilchenschule Obertrubach-Bärnfels, 09.10.2010

Radiosignals of extensive air showers

Radio signal of EAS

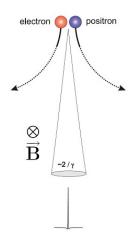
- Mainly synchtroton radiation of electrons and positrons
- Emitted into a narrow cone into particle's direction of travel
- Short bipolar radio pulse
 10 100 MHz range
 Duration shorter than 10 ns
- Proof of detection principle by LOPES



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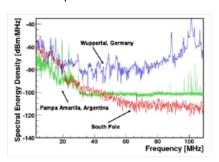
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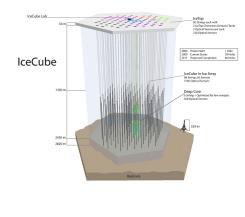


Advantages of the South Pole

Local advantages

- Low radio noise level
- Good weather conditions
- Existing experiments: IceCube, IceTop





Synergy effects

Neutrino search

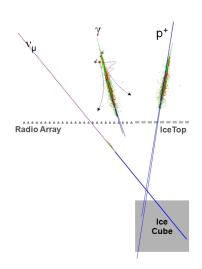
Increase surface veto area for atmospheric neutrinos (main background)

Primary particle composition

Use different handles from three different detectors (muons, electrons, shower maximum)

PeV Gamma Ray Search

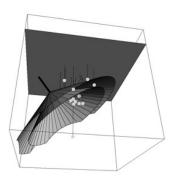
Search for showers with low muon numbers (inverted veto)



RICE

RICE

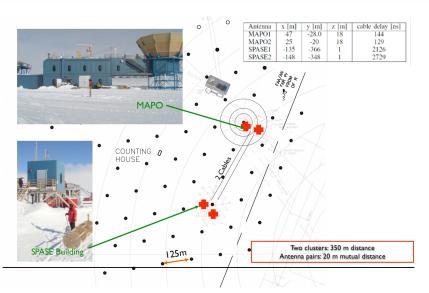
- Radio Ice Cherenkov Experiment
- 18 radio antennas in a depth between 100 and 300m
- Neutrino detection using the Askaryan effect



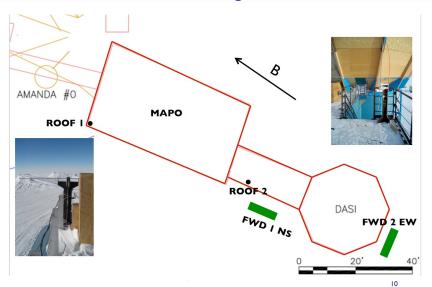
RICE Trigger

- Inlce
- External
- Veto
- Surface

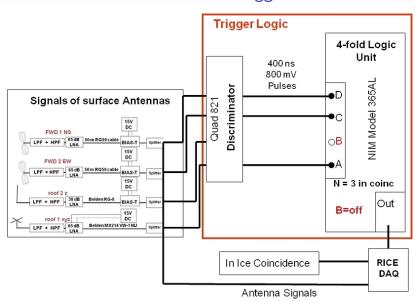
RICE Surface Configuration 2009



RICE Surface Configuration 2010



RICE Surface Trigger

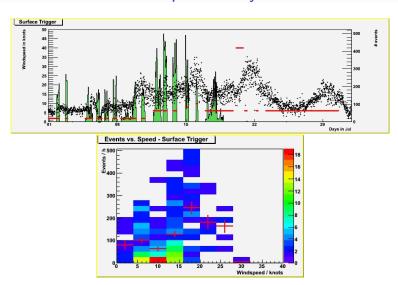


Diploma thesis plans

Search for coincidences in IceTop and RICE Surface Antennas

- Data quality
 - Wind dependence
 - Pressure dependence
 - Daytime dependence
 - Periodic events
- Analysis of signal waveforms and spectra
- Search for coincident events

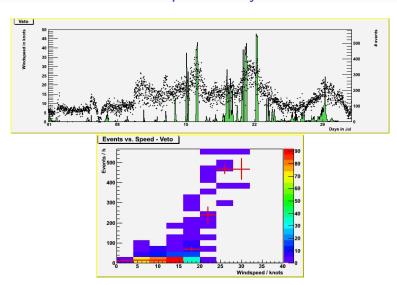
Windspeed Analysis



First analysis: Surface trigger not correlated to wind



Windspeed Analysis



First analysis: Veto Trigger clearly correlateted, machanism unknown

